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### Collinear factorisation for $e^+e^-$ collisions

Based on: 1909.03886 (SF), 1911.12040 (Bertone, Cacciari, SF, Stagnitto) 2105.06688 (SF), 2108.10261 (SF, Mattelaer, Zaro, Zhao) 2207.03265 (Bertone, Cacciari, SF, Stagnitto, Zaro, Zhao)

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Assumption:

Somewhere, someone will build an  $e^+e^-$  collider (linear or circular)

### Cross sections stemming from e<sup>+</sup>e<sup>-</sup> collisions are plagued by large logs that must be resummed

One way to do that is by means of collinear factorisation; another, with YFS

Either way, the so-called precision tools currently available are not sufficiently accurate to meet the expected precision targets – only limited progress made since the 90's Consider a generic cross section, sufficiently inclusive:

$$\sigma = \alpha^b \sum_{n=0}^{\infty} \alpha^n \sum_{i=0}^n \sum_{j=0}^n \varsigma_{n,i,j} L^i \ell^j$$

This is symbolic, and only useful to expose the presence of:

$$\ell = \log \frac{Q^2}{\langle E_\gamma \rangle^2}, \qquad L = \log \frac{Q^2}{m^2}$$

Numerology: consider the production of  $Z \rightarrow ll$  at:

• 
$$\sqrt{Q^2} = m_Z$$

$$L = 24.18 \implies \frac{\alpha}{\pi}L = 0.06$$
$$0 \le m_{ll} \le m_Z, \quad \ell = 8.29 \implies \frac{\alpha}{\pi}\ell = 0.02$$
$$m_Z - 1 \text{ GeV} \le m_{ll} \le m_Z, \quad \ell = 13.66 \implies \frac{\alpha}{\pi}\ell = 0.034$$

Consider a generic cross section, sufficiently inclusive:

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$$\ell = \log \frac{Q^2}{\langle E_\gamma \rangle^2}, \qquad L = \log \frac{Q^2}{m^2}$$

Numerology: consider the production of  $Z \rightarrow ll$  at:

• 
$$\sqrt{Q^2} = 500 \text{ GeV}$$

$$L = 24.59 \implies \frac{\alpha}{\pi}L = 0.068$$
$$0 \le m_{ll} \le m_Z, \quad \ell = 1.46 \implies \frac{\alpha}{\pi}\ell = 0.0036$$
$$m_Z - 1 \text{ GeV} \le m_{ll} \le m_Z, \quad \ell = 4.51 \implies \frac{\alpha}{\pi}\ell = 0.01$$

It takes a lot of brute force (i.e. fixed-order results to some  $\mathcal{O}(\alpha^n)$ ) to overcome the enhancements due to L and  $\ell$ .

It is always convenient to first improve by means of factorisation formulae:

$$d\sigma(L,\ell) = \mathcal{K}_{soft}(\ell;L)\beta(L)d\mu \tag{1}$$

$$= \mathcal{K}_{coll}(L;\ell) \otimes d\hat{\sigma}(\ell) \tag{2}$$

Use of:

- (1) YFS (resummation of  $\ell$ )
- (2) collinear factorisation (resummation of L)

Common features:  $\mathcal{K}$  is an *all-order* universal factor;  $\beta$  and  $d\hat{\sigma}$  are process-specific and computed order by order (still brute force, but one needs comparatively less)

### YFS

Aim: soft resummation for:

$$\left\{e^+(p_1) + e^-(p_2) \longrightarrow X(p_X) + \sum_{i=0}^n \gamma(k_n)\right\}_{n=0}^{\infty}$$

Achieved with:

$$d\sigma(L,\ell) = \mathcal{K}_{soft}(\ell;L)\beta(L)d\mu$$
  
=  $e^{Y(p_1,p_2,p_X)}\sum_{n=0}^{\infty}\beta_n \left(\mathcal{R}p_1,\mathcal{R}p_2,\mathcal{R}p_X;\{k_i\}_{i=0}^n\right)d\mu_{X+n\gamma}$ 

This is symbolic, and stands for both the EEX and CEEX approaches [hep-ph/0006359 Jadach, Ward, Was] that build upon the original YFS work [Ann.Phys.13(61)379]

EEX: exclusive (in the photons) exponentiation, matrix element level

CEEX: coherent exclusive (in the photons) exponentiation, amplitude level, including interference

### YFS

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Achieved with:

$$d\sigma(L,\ell) = e^{Y(p_1,p_2,p_X)} \sum_{n=0}^{\infty} \beta_n \left(\mathcal{R}p_1, \mathcal{R}p_2, \mathcal{R}p_X; \{k_i\}_{i=0}^n\right) d\mu_{X+n\gamma}$$

- Y essentially universal (process dependence only through kinematics); resums  $\ell$
- The soft-finite  $\beta_n$  are process-specific, and are constructed by means of local subtractions involving matrix elements and eikonals (i.e. *not* BN)

$$\beta_n = \alpha^b \sum_{i=0}^n \alpha^i \sum_{j=0}^i c_{n,i,j} L^j$$

• For a given n, matrix elements have different multiplicities, hence the need for the kinematic mapping  ${\cal R}$ 

## Collinear factorisation

Aim: collinear resummation for:

$$\left\{k(p_k) + l(p_l) \longrightarrow X(p_X) + \sum_{i=0}^n a_i(k_n)\right\}_{n=0}^{\infty} \qquad a_i = e^{\pm}, \gamma \dots$$

with initial-state particles stemming from beams:

$$(k,l) = (e^+, e^-), \quad (k,l) = (e^+, \gamma), \quad (k,l) = (\gamma, e^-), \quad (k,l) = (\gamma, \gamma), \dots$$

Master formula:

$$d\sigma(L,\ell) = \mathcal{K}_{coll}(L;\ell) \otimes d\hat{\sigma}(\ell)$$
  
$$\longrightarrow d\sigma_{kl} = \sum_{ij} \int dz_{+} dz_{-} \Gamma_{i/k}(z_{+},\mu^{2},m^{2}) \Gamma_{j/l}(z_{-},\mu^{2},m^{2})$$

$$\times d\hat{\sigma}_{ij}(z_+p_k, z_-p_l, \mu^2; p_X, \{k_i\}_{i=0}^n)$$

- $\Gamma_{\alpha/\beta}$  universal (the PDF); resums L
- The collinear-finite  $d\hat{\sigma}_{ij}$  are process-specific, and are the standard short-distance matrix elements, constructed order by order (*with* BN). May or may not include resummation of other large logs (including  $\ell$ )

## YFS vs collinear factorisation

Both are systematically improvable in perturbation theory: in YFS the  $\beta_n$ 's (fixed-order), in collinear factorisation both the PDFs (logarithmic accuracy) and the  $d\hat{\sigma}$ 's (fixed-order, resummation)

- + YFS: very little room for systematics. Exceptions are the kinematic mapping  $\mathcal{R}$ , and the quark masses (when the quarks are radiators). Renormalisation schemes??
- Collinear factorisation: systematic variations much larger. At the LL (used in phenomenology so far) a rigorous definition of uncertainties is impossible (parameters are arbitrary), and comparisons with YFS are largely fine tuned
- YFS: the computations of  $\beta_n$  are not standard (EEX) and highly non-trivial (CEEX)
- + Collinear factorisation: the computations of  $d\hat{\sigma}_{ij}$  are standard

## Collinear factorisation



### $d\sigma = \mathsf{PDF} \star \mathsf{PDF} \star d\hat{\sigma}$

### PDFs collect (universal) small-angle dynamics

*Very* similar to  $QCD^*$ , with some notable differences:

- PDFs and power-suppressed terms can be computed perturbatively
- An object (e.g.  $e^-$ ) may play the role of both particle and parton

As in QCD, a particle is a physical object, a parton is not

\*Side benefit: we can recycle a lot of what is done for LHC

All physics simulations based on collinear factorisation done so far are based on a LL-accurate picture

This is not tenable at high energies/high statistics:

- accuracy is insufficient (see e.g.  $W^+W^-$  production)
- systematics not well defined (e.g.  $\alpha$  is literally an arbitrary parameter)

Bear in mind:

There is no precision physics without the ability of assessing uncertainties All physics simulations based on collinear factorisation done so far are based on a LL-accurate picture

This is not tenable at high energies/high statistics:

- $\blacklozenge$  accuracy is insufficient (see e.g.  $W^+W^-$  production)
- systematics not well defined (e.g.  $\alpha$  is literally an arbitrary parameter)

Step 0 was to upgrade PDFs from LL to NLL accuracy: increase of precision, and meaningful systematics, in particular factorisation-scheme dependence

z-space LO+LL PDFs  $(\alpha \log(Q^2/m^2))^k$ :

 $\sim 1992$ 

- $\blacktriangleright$   $0 \leq k \leq \infty$  for  $z \simeq 1$  (Gribov, Lipatov)
- $\blacktriangleright$   $0 \leq k \leq 3$  for z < 1 (Skrzypek, Jadach; Cacciari, Deandrea, Montagna, Nicrosini; Skrzypek)
- matching between these two regimes
- ▶ for  $e^-$

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- $\blacktriangleright$  for  $e^-$

*z*-space NLO+NLL PDFs  $(\alpha \log(Q^2/m^2))^k + \alpha (\alpha \log(Q^2/m^2))^{k-1}$ :

 $\longrightarrow 1909.03886, 1911.12040, 2105.06688 \quad (\mathsf{Bertone}, \mathsf{Cacciari}, \mathsf{Frixione}, \mathsf{Stagnitto})$ 

- ▶  $0 \le k \le \infty$  for  $z \simeq 1$
- ▶  $0 \le k \le 3$  for  $z < 1 \iff \mathcal{O}(\alpha^3)$
- matching between these two regimes
- ▶ for  $e^+$ ,  $e^-$ , and  $\gamma$
- both numerical and analytical
- $\blacktriangleright$  factorisation schemes:  $\overline{\mathrm{MS}}$  and  $\Delta$  (that has DIS-like features)

NLO initial conditions (1909.03886) Conventions for the perturbative coefficients:

$$\Gamma_i = \Gamma_i^{[0]} + \frac{\alpha}{2\pi} \Gamma_i^{[1]} + \mathcal{O}(\alpha^2)$$

Results:

$$\begin{split} &\Gamma_i^{[0]}(z,\mu_0^2) &= \delta_{ie^-}\delta(1-z) \\ &\Gamma_{e^-}^{[1]}(z,\mu_0^2) &= \left[\frac{1+z^2}{1-z}\left(\log\frac{\mu_0^2}{m^2}-2\log(1-z)-1\right)\right]_+ + K_{ee}(z) \\ &\Gamma_{\gamma}^{[1]}(z,\mu_0^2) &= \frac{1+(1-z)^2}{z}\left(\log\frac{\mu_0^2}{m^2}-2\log z-1\right) + K_{\gamma e}(z) \\ &\Gamma_{e^+}^{[1]}(z,\mu_0^2) &= 0 \end{split}$$

Note:

▶ Meaningful only if  $\mu_0 \sim m$ 

▶ In  $\overline{\text{MS}}$ ,  $K_{ij}(z) = 0$ ; in general, these functions define a factorisation scheme

Bear in mind that PDFs are fully defined only after adopting a definite *factorisation scheme*, which is the choice of the finite terms associated with the subtraction of the collinear poles

(done by means of the  $K_{ij}(z)$  functions)

- $\blacklozenge$  1911.12040  $\longrightarrow \overline{\mathrm{MS}}$
- 2105.06688  $\longrightarrow$  a DIS-like scheme (called  $\Delta$ )

At variance with the QCD case, there is also an interesting *renormalisation-scheme* dependence of QED PDFs (not discussed in this talk)

### Definition of the $\Delta$ scheme

The idea: minimise the impact of NLO corrections

$$K_{ee}^{(\Delta)}(z) = \left[\frac{1+z^2}{1-z}\left(2\log(1-z)+1\right)\right]_+$$
$$K_{\gamma e}^{(\Delta)}(z) = \frac{1+(1-z)^2}{z}\left(2\log z+1\right)$$

Thus:

$$\begin{split} \Gamma_{e^{-}}(z,\mu_{0}^{2}) &= \delta(1-z) + \frac{\alpha}{2\pi} \left[ \frac{1+z^{2}}{1-z} \right]_{+} \log \frac{\mu_{0}^{2}}{m^{2}} \\ \Gamma_{\gamma}(z,\mu_{0}^{2}) &= \frac{\alpha}{2\pi} \frac{1+(1-z)^{2}}{z} \log \frac{\mu_{0}^{2}}{m^{2}} \end{split}$$

This is exactly the same as at the LO when  $\mu_0 = m$ (a convenient choice for several reasons)

▶ The kernels  $K_{ij}(z)$  also enter the evolution equations (see later)

# NLL evolution (1911.12040, 2105.06688)

General idea: solve the evolution equations starting from the initial conditions computed previously

$$\frac{\partial\Gamma_i(z,\mu^2)}{\partial\log\mu^2} = \frac{\alpha(\mu)}{2\pi} \left[P_{ij}\otimes\Gamma_j\right](z,\mu^2) \iff \frac{\partial\Gamma(z,\mu^2)}{\partial\log\mu^2} = \frac{\alpha(\mu)}{2\pi} \left[\mathbb{P}\otimes\Gamma\right](z,\mu^2),$$

#### Two ways:

- Mellin space: suited to both numerical solution and all-order, large-z analytical solution (called *asymptotic solution*). <u>Dominant</u>
- Directly in z space in an integrated form: suited to fixed-order, all-z analytical solution (called *recursive solution*). Subleading

Formulae are simpler when working with a single lepton family; the physical picture is the same as that with multiple families A technicality: owing to the running of  $\alpha$ , it is best to evolve in t rather than in  $\mu$ , with: (~ Furmanski, Petronzio)

$$t = \frac{1}{2\pi b_0} \log \frac{\alpha(\mu)}{\alpha(\mu_0)}$$
  
=  $\frac{\alpha(\mu)}{2\pi} L - \frac{\alpha^2(\mu)}{4\pi} \left( b_0 L^2 - \frac{2b_1}{b_0} L \right) + \mathcal{O}(\alpha^3), \qquad L = \log \frac{\mu^2}{\mu_0^2}$ 

#### Note:

- $\blacktriangleright$  t  $\longleftrightarrow$   $\mu$ ; notation-wise, the dependence on t is equivalent to the dependence on  $\mu$
- $\blacktriangleright t = 0 \iff \mu = \mu_0$
- ► L is my "large log"
- Fricky: fixed- $\alpha$  expressions are obtained with  $t = \alpha L/(2\pi)$  (and not t = 0)

# Mellin space

Introduce the evolution operator  $\mathbb{E}_N$ 

 $\Gamma_N(\mu^2) = \mathbb{E}_N(t) \,\Gamma_{0,N} \,, \qquad \mathbb{E}_N(0) = I \,, \qquad \Gamma_{0,N} \equiv \Gamma_N(\mu_0^2)$ 

The PDFs evolution equations are then re-expressed by means of an evolution equation for the evolution operator:

$$\frac{\partial \mathbb{E}_{N}^{(K)}(t)}{\partial t} = b_{0}\alpha(\mu)\mathbb{K}_{N}\left(I + \frac{\alpha(\mu)}{2\pi}\mathbb{K}_{N}\right)^{-1}\mathbb{E}_{N}^{(K)}(t) + \frac{b_{0}\alpha^{2}(\mu)}{\beta(\alpha(\mu))}\sum_{k=0}^{\infty}\left(\frac{\alpha(\mu)}{2\pi}\right)^{k} \times \left(I + \frac{\alpha(\mu)}{2\pi}\mathbb{K}_{N}\right)\mathbb{P}_{N}^{[k]}\left(I + \frac{\alpha(\mu)}{2\pi}\mathbb{K}_{N}\right)^{-1}\mathbb{E}_{N}^{(K)}(t)$$

Can be solved numerically

Can be solved analytically in a closed form under simplifying assumptions. Chiefly: large-z is equivalent to large-N

# Asymptotic $\overline{\mathrm{MS}}$ solution

Non-singlet  $\equiv$  singlet; photon is more complicated

$$\Gamma_{\rm NLL}(z,\mu^2) \xrightarrow{z \to 1} \frac{e^{-\gamma_{\rm E}\xi_1} e^{\hat{\xi}_1}}{\Gamma(1+\xi_1)} \xi_1 (1-z)^{-1+\xi_1} \\ \times \left\{ 1 + \frac{\alpha(\mu_0)}{\pi} \left[ \left( L_0 - 1 \right) \left( A(\xi_1) + \frac{3}{4} \right) - 2B(\xi_1) + \frac{7}{4} \right. \\ \left. + \left( L_0 - 1 - 2A(\xi_1) \right) \log(1-z) - \log^2(1-z) \right] \right\}$$

where  $L_0 = \log \mu_0^2 / m^2$ , and:

$$A(\kappa) = -\gamma_{\rm E} - \psi_0(\kappa)$$
  
$$B(\kappa) = \frac{1}{2}\gamma_{\rm E}^2 + \frac{\pi^2}{12} + \gamma_{\rm E}\psi_0(\kappa) + \frac{1}{2}\psi_0(\kappa)^2 - \frac{1}{2}\psi_1(\kappa)$$

with:

$$\begin{aligned} \xi_1 &= 2t - \frac{\alpha(\mu)}{4\pi^2 b_0} \left( 1 - e^{-2\pi b_0 t} \right) \left( \frac{20}{9} n_F + \frac{4\pi b_1}{b_0} \right) \\ &= 2t + \mathcal{O}(\alpha t) = \eta_0 + \dots \\ \hat{\xi}_1 &= \frac{3}{2} t + \frac{\alpha(\mu)}{4\pi^2 b_0} \left( 1 - e^{-2\pi b_0 t} \right) \left( \lambda_1 - \frac{3\pi b_1}{b_0} \right) \\ &= \frac{3}{2} t + \mathcal{O}(\alpha t) = \lambda_0 \eta_0 + \dots \\ \lambda_1 &= \frac{3}{8} - \frac{\pi^2}{2} + 6\zeta_3 - \frac{n_F}{18} (3 + 4\pi^2) \end{aligned}$$

Remember that:

$$t = \frac{1}{2\pi b_0} \log \frac{\alpha(\mu)}{\alpha(\mu_0)}$$
  
=  $\frac{\alpha(\mu)}{2\pi} L - \frac{\alpha^2(\mu)}{4\pi} \left( b_0 L^2 - \frac{2b_1}{b_0} L \right) + \mathcal{O}(\alpha^3), \qquad L = \log \frac{\mu^2}{\mu_0^2}$ 

.

### Asymptotic $\Delta$ solution

Non-singlet  $\equiv$  singlet; photon is trivial

$$\Gamma_{\rm NLL}(z,\mu^2) \xrightarrow{z \to 1} \frac{e^{-\gamma_{\rm E}\xi_1} e^{\hat{\xi}_1}}{\Gamma(1+\xi_1)} \xi_1 (1-z)^{-1+\xi_1} \\ \times \left[ \left( 1 + \frac{3\alpha(\mu_0)}{4\pi} L_0 \right) \sum_{p=0}^{\infty} S_{1,p}(z) - \frac{\alpha(\mu_0)}{\pi} L_0 \sum_{p=0}^{\infty} S_{2,p}(z) \right]$$

The  $\mathcal{S}_{i,p}(z)$  functions are increasingly suppressed at  $z \to 1$  with growing p. The dominant behaviour is:

$$\Gamma_{\rm NLL}(z,\mu^2) \xrightarrow{z \to 1} \frac{e^{-\gamma_{\rm E}\xi_1} e^{\hat{\xi}_1}}{\Gamma(1+\xi_1)} \xi_1 (1-z)^{-1+\xi_1} \\
\times \left[ \frac{\alpha(\mu)}{\alpha(\mu_0)} + \frac{\alpha(\mu)}{\pi} L_0 \left( A(\xi_1) + \log(1-z) + \frac{3}{4} \right) \right]$$

A vastly different logarithmic behaviour w.r.t. the  $\overline{MS}$  case However,  $\Gamma_{NLL}^{(\overline{MS})} - \Gamma_{NLL}^{(\Delta)} = \mathcal{O}(\alpha^2)$  Key facts

• Both  $\overline{\text{MS}}$  and  $\Delta$  results feature an integrable singularity at  $z \to 1$ , basically identical to the LL one

### Key facts

- Both  $\overline{\text{MS}}$  and  $\Delta$  results feature an integrable singularity at  $z \to 1$ , basically identical to the LL one
- $\blacklozenge$  In addition to that, in  $\overline{\mathrm{MS}}$  there are single and double logarithmic terms

#### Key facts

- Both  $\overline{\text{MS}}$  and  $\Delta$  results feature an integrable singularity at  $z \to 1$ , basically identical to the LL one
- $\blacklozenge$  In addition to that, in  $\overline{\mathrm{MS}}$  there are single and double logarithmic terms
- Owing to the integrable singularity, it is essential to have large-z analytical results: the PDFs convoluted with cross sections are obtained by matching the small- and intermediate-z numerical solution with the large-z analytical one

Analytical recursive solutions are used as cross-checks

#### A look at the photon:

$$\begin{split} \Gamma_{\gamma}^{(\overline{\mathrm{MS}})}(z,\mu^{2}) & \xrightarrow{z \to 1} & \frac{t\alpha(\mu_{0})^{2}}{\alpha(\mu)} \frac{3}{2\pi\xi_{1}} \log(1-z) - \frac{t\alpha(\mu_{0})^{3}}{\alpha(\mu)} \frac{1}{2\pi^{2}\xi_{1}} \log^{3}(1-z) \\ \Gamma_{\gamma}^{(\Delta)}(z,\mu^{2}) & \xrightarrow{z \to 1} & \frac{1}{2\pi} \frac{\alpha^{2}(\mu_{0})}{\alpha(\mu)} \frac{1+(1-z)^{2}}{z} L_{0} + \frac{1}{2\pi\xi_{1}} \frac{t\alpha^{2}(\mu_{0})}{\alpha(\mu)} L_{0} \\ & -\frac{t\alpha(\mu)}{2\pi\xi_{1}} \frac{e^{-\gamma_{\mathrm{E}}\xi_{1}}e^{\hat{\xi}_{1}}}{\Gamma(1+\xi_{1})} (1-z)^{\xi_{1}} L_{0} \,. \end{split}$$

■  $\overline{\text{MS}}$  vs  $\Delta$  exhibits the same pattern as for (non-)singlet: logarithmic terms dominate at  $z \rightarrow 1$  in  $\overline{\text{MS}}$ , but are absent in  $\Delta$ 

All physics simulations based on collinear factorisation done so far are based on a LL-accurate picture

This is not tenable at high energies/high statistics:

 $\blacklozenge$  accuracy is insufficient (see e.g.  $W^+W^-$  production)

systematics not well defined

Step 0 was to upgrade PDFs from LL to NLL accuracy

Step 1 (2207.03265, Bertone, Cacciari, Frixione, Stagnitto, Zaro, Zhao) is to include in the NLL PDFs thus obtained the ingredients necessary for sensible phenomenology, in particular:

evolution with all fermion families (leptons and quarks), including their respective mass thresholds

▶ renormalisation schemes:  $\overline{\text{MS}}$ ,  $\alpha(m_Z)$ , and  $G_{\mu}$ 

assess implications by studying realistic observables in physical processes

#### Sample results for:

$$e^{+}e^{-} \longrightarrow q\bar{q}$$

$$e^{+}e^{-} \longrightarrow t\bar{t}$$

$$e^{+}e^{-} \longrightarrow W^{+}W^{-}$$

with  $q\bar{q}$  production (massless quarks) restricted to ISR QED radiation. The other two are in the SM

NLO accuracy, automated generation with MG5\_aMC@NLO (this version now public, but not yet as v3.X)

What is plotted:

$$\sigma(\tau_{min}) = \int d\sigma \,\Theta\left(\tau_{min} \le \frac{M_{p\bar{p}}^2}{s}\right) \,, \qquad p = q \,, t \,, W^+$$

 $\tau_{min} \sim 1$  is sensitive to soft emissions (not resummed)

### Dependence of PDFs on factorisation scheme



Very large dependence at the NLL at  $z \to 1$  ( $\mathcal{O}(1)$ ); this is particularly significant (but unphysical!) since the electron has an integrable divergence there

Electron at NLL in the Delta scheme close to the LL result (differences of  $\mathcal{O}(5\%)$ )

### Dependence of observables on factorisation scheme



 $\mathcal{O}(1)$  differences for PDFs down to  $\mathcal{O}(10^{-4} - 10^{-3})$  for observables

In the  $\overline{\mathrm{MS}}$  scheme, huge cancellations between PDFs and short-distance cross sections Behaviour qualitatively similar for different renormalisation schemes

### Factorisation vs renormalisation scheme dependence



Renormalisation-scheme dependence much larger than factorisation-scheme dependence, with process-dependent pattern

Depending on the precision, renormalisation scheme is an informed choice; factorisation scheme always induces a systematic

### NLL vs LL



#### Effects are non trivial

Pattern dependent on the process (and on the observable) as well as on the renormalisation scheme

# Impact of $\gamma\gamma$ channel



Essentially independent of factorisation and renormalisation schemes: a genuine physical effect

Utterly negligible for  $t\bar{t}$ , significant for  $W^+W^-$  – process dependence is not surprising

# Thus:

- ► The inclusion of NLL contributions into the electron PDF has an impact of O(1%) (precise figures are observable and renormalisation-scheme dependent)
- This estimate does not include the effects of the photon PDF
- ► The comparison between  $\overline{\mathrm{MS}}$  and  $\Delta$ -based results shows differences compatible with non-zero  $\mathcal{O}(\alpha^2)$  effects, as expected
- ▶ Renormalisation-scheme dependence is of  $\mathcal{O}(0.5\%)$

If the target is a  $10^{-\rm some\ large\ number}$  relative precision, these effects must be taken into account

# Outlook

- Increasing the precision of theoretical results will be essential for the success of the physics programs at future  $e^+e^-$  colliders
- QED collinear factorisation is very similar to its QCD counterpart: it is possible to recycle many of the techniques invented for the LHC
- QED PDFs are now NLL accurate. It is important that they be used in the context of realistic simulations
- LEP-era YFS-based results are not up to the task. What to do is in principle clear; in practice, it looks painful and of limited scope (count the number of applications)

**EXTRA SLIDES** 

#### z space

Use integrated PDFs (so as to simplify the treatment of endpoints)

$$\mathcal{F}(z,t) = \int_0^1 dy \,\Theta(y-z)\,\Gamma(y,\mu^2) \quad \Longrightarrow \quad \Gamma(z,\mu^2) = -\frac{\partial}{\partial z}\mathcal{F}(z,t)$$

in terms of which the formal solution of the evolution equation is:

$$\mathcal{F}(z,t) = \mathcal{F}(z,0) + \int_0^t du \, \frac{b_0 \alpha^2(u)}{\beta(\alpha(u))} \left[\mathbb{P} \,\overline{\otimes} \,\mathcal{F}\right](z,u)$$

By inserting the representation:

$$\mathcal{F}(z,t) = \sum_{k=0}^{\infty} \frac{t^k}{k!} \left( \mathcal{J}_k^{\text{LL}}(z) + \frac{\alpha(t)}{2\pi} \, \mathcal{J}_k^{\text{NLL}}(z) \right)$$

on both sides of the solution, one obtains recursive equations, whereby a  $\mathcal{J}_k$  is determined by all  $\mathcal{J}_p$  with p < k. The recursion starts from  $\mathcal{J}_0$ , which are the integrated initial conditions

For the record, the recursive equations are:

$$\begin{aligned}
\mathcal{J}_{k}^{\text{LL}} &= \mathbb{P}^{[0]} \overline{\otimes} \mathcal{J}_{k-1}^{\text{LL}} \\
\mathcal{J}_{k}^{\text{NLL}} &= (-)^{k} (2\pi b_{0})^{k} \mathcal{F}^{[1]}(\mu_{0}^{2}) \\
&+ \sum_{p=0}^{k-1} (-)^{p} (2\pi b_{0})^{p} \left( \mathbb{P}^{[0]} \overline{\otimes} \mathcal{J}_{k-1-p}^{\text{NLL}} + \mathbb{P}^{[1]} \overline{\otimes} \mathcal{J}_{k-1-p}^{\text{LL}} \\
&- \frac{2\pi b_{1}}{b_{0}} \mathbb{P}^{[0]} \overline{\otimes} \mathcal{J}_{k-1-p}^{\text{LL}} \right)
\end{aligned}$$

We have computed these for  $k \leq 3$  ( $\mathcal{J}^{LL}$ ) and  $k \leq 2$  ( $\mathcal{J}^{NLL}$ ), ie to  $\mathcal{O}(\alpha^3)$ Results in 1911.12040 and its ancillary files

### Large-z singlet and photon

As for the non-singlet, start from the asymptotic AP kernel expressions:

$$\mathbb{P}_{\mathrm{S},N} \xrightarrow{N \to \infty} \begin{pmatrix} -2\log\bar{N} + 2\lambda_0 & 0\\ 0 & -\frac{2}{3}n_F \end{pmatrix} + \frac{\alpha}{2\pi} \begin{pmatrix} \frac{20}{9}n_F\log\bar{N} + \lambda_1 & 0\\ 0 & -n_F \end{pmatrix} + \mathcal{O}(1/N) + \mathcal{O}(\alpha^2)$$

This implies

$$(\mathbb{E}_N)_{SS} = E_N$$
$$M^{-1} [(\mathbb{E}_N)_{\gamma\gamma}] = \frac{\alpha(\mu_0)}{\alpha(\mu)} \,\delta(1-z)$$

- $\Rightarrow$  Singlet  $\equiv$  non-singlet
  - Photon  $\equiv$  initial condition +  $\alpha(0)$  scheme

Photon  $\equiv$  initial condition  $+ \alpha(0)$  scheme  $\Longrightarrow$  $\Gamma_{\gamma}(z,\mu^2) = \frac{1}{2\pi} \frac{\alpha(\mu_0)^2}{\alpha(\mu)} \frac{1 + (1-z)^2}{z} \left( \log \frac{\mu_0^2}{m^2} - 2\log z - 1 \right).$ 

Or:  $\sim$  Weizsaecker-Williams function, plus the natural emergence of a small scale in the argument of  $\alpha$ 

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 $\rightarrow 1/N$  suppression of off-diagonal terms in the evolution operator is over-compensated by the  $\delta$ -like peak of the electron initial-condition Photon  $\equiv$  initial condition  $+ \alpha(0)$  scheme  $\Longrightarrow$ 

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By solving the  $2 \times 2$  system e.g. in MS:

$$\Gamma_{\gamma}^{(\overline{\mathrm{MS}})}(z,\mu^2) \xrightarrow{z \to 1} \frac{t\alpha(\mu_0)^2}{\alpha(\mu)} \frac{3}{2\pi\xi_1} \log(1-z) - \frac{t\alpha(\mu_0)^3}{\alpha(\mu)} \frac{1}{2\pi^2\xi_1} \log^3(1-z)$$

## A remarkable fact

Our asymptotic solutions, expanded in  $\alpha$ , feature **all** of the terms:

$$\frac{\log^q (1-z)}{1-z} \qquad \text{singlet, non-singlet} \\ \log^q (1-z) \qquad \text{photon}$$

of our recursive solutions

Non-trivial; stems from keeping subleading terms (at  $z \rightarrow 1$ ) in the AP kernels