

# Kaluza-Klein Dark Matter

-

## Direct Detection vis-a-vis LHC

Sebastian Arrenberg  
University of Zürich

Based on hep-ph/0805.4210 by  
Sebastian Arrenberg  
Laura Baudis  
Kyoungchul Kong  
Konstantin T. Matchev  
Jonghee Yoo

(approved to appear in PRD)

IDM 2008 – Identification of Dark Matter 2008  
18.08.2008  
Stockholm

Short overview of Universal Extra Dimensions (UEDs)

The relic density of Kaluza-Klein particles

Elastic scattering cross sections – predictions and limits

Limits on the degeneracy parameter  $\Delta$ , the Higgs mass  $m_h$  and spin-dependent WIMP-nucleon couplings

- all Standard Model particles are promoted to one or more compactified flat extra dimensions

- infinite number of new particles (Kaluza-Klein tower)

- tree level masses:  $m_n^2 = m^2 + \frac{n^2}{R^2}$ 
  - ← quantum number labelling the  $n^{\text{th}}$  KK mode
  - mass of the associated SM particle
  - compactification scale

- high degree of mass degeneracy

→ radiative corrections are of crucial importance

- including radiative corrections yields KK parity  $(-1)^n$  conservation

→ stable level 1 particles

→ possible dark matter candidates

- WIMP candidates:

		5D			6D
<del><math>G_1</math></del>	<del><math>\nu_1</math></del>	$\gamma_1$	$Z_1$	$H_1$	$\gamma_H$
					$Z_H$

high degree of mass degeneracy

—► coannihilations with all  $n=1$  KK particles were taken into account

lightest particle obeys the Boltzmann equation

$$\frac{dn}{dt} = -3 Hn - \langle \sigma_{eff} v \rangle (n^2 - n_{eq}^2)$$

with

$$\sigma_{eff}(x) = \sum_{ij} \sigma_{ij} \frac{g_i g_j}{g_{eff}^2} (1 + \Delta_i)^{\frac{3}{2}} (1 + \Delta_j)^{\frac{3}{2}} e^{-x(\Delta_i + \Delta_j)} \quad x = \frac{m_1}{T}$$

$$g_{eff}(x) = \sum_i g_i (1 + \Delta_i)^{\frac{3}{2}} e^{-x \Delta_i} \quad \Delta_i = \frac{m_i - m_1}{m_1}$$

mass degeneracy  
parameter

## What about the masses?

assume vanishing boundary interactions at the cut-off scale (minimal UED)

—► radiative corrections to the masses can be computed (hep-ph/0204342)

LKP using MUED framework

$\swarrow$   
 $\searrow$

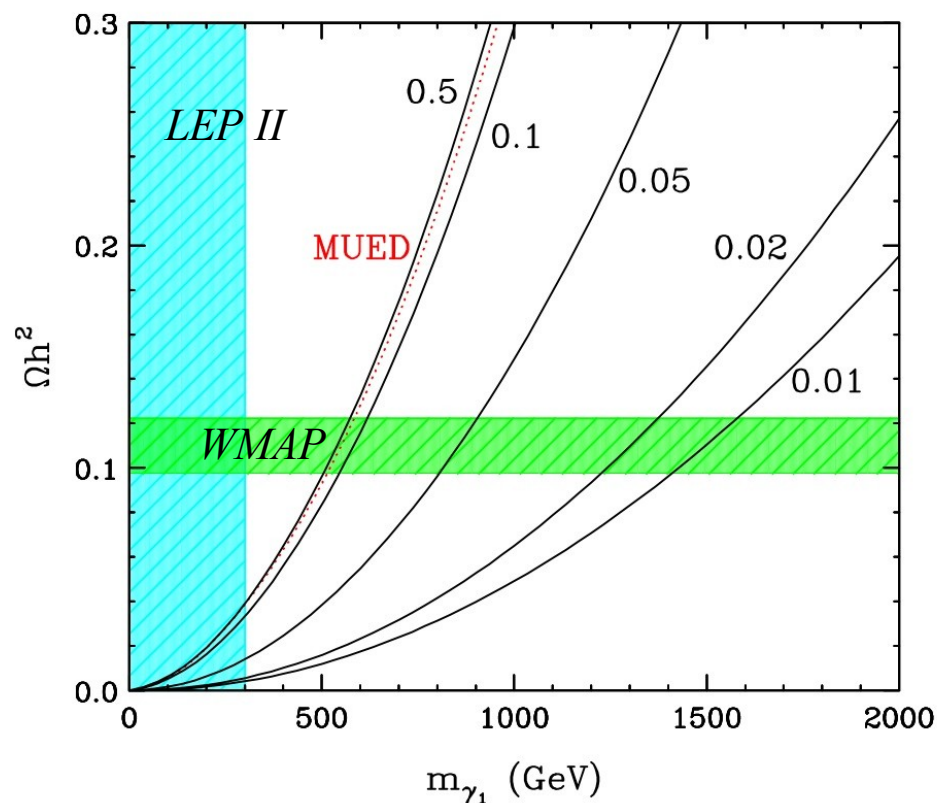
5D  $\gamma_1$

6D  $\gamma_H$

But consider other possibilities as well...

$\gamma_1$

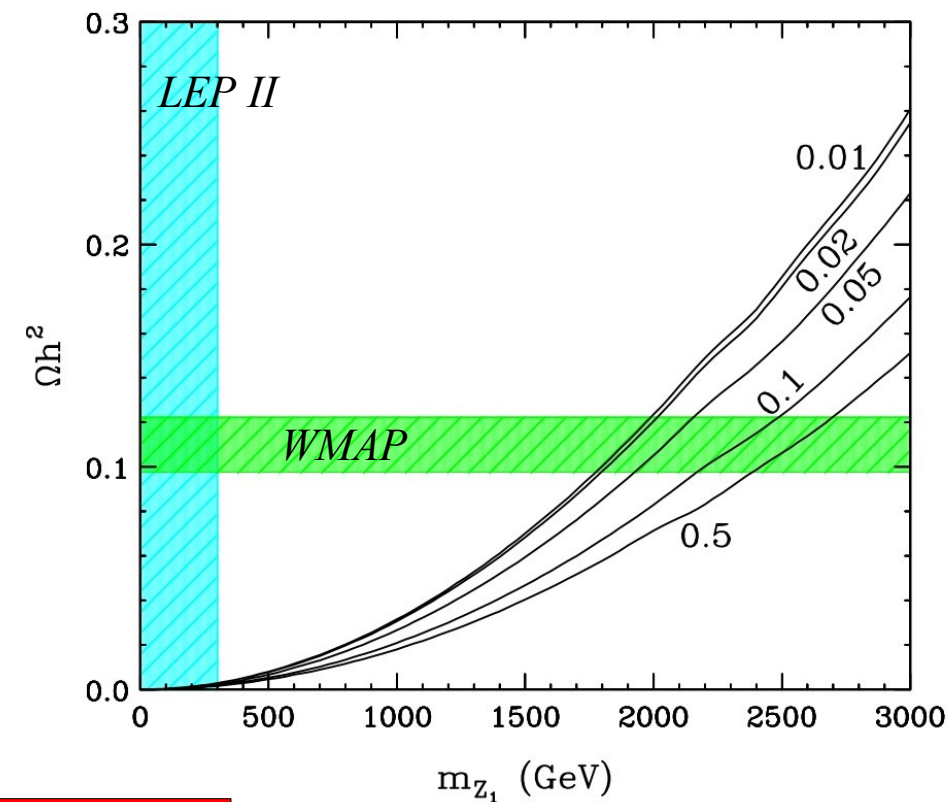
- 1) MUED framework
- 2) - assume certain mass splitting  $\Delta$  between LKP and KK quarks
  - fix rest of the spectrum using MUED



- coannihilations are indeed important
- the sign of the effect cannot easily be predicted

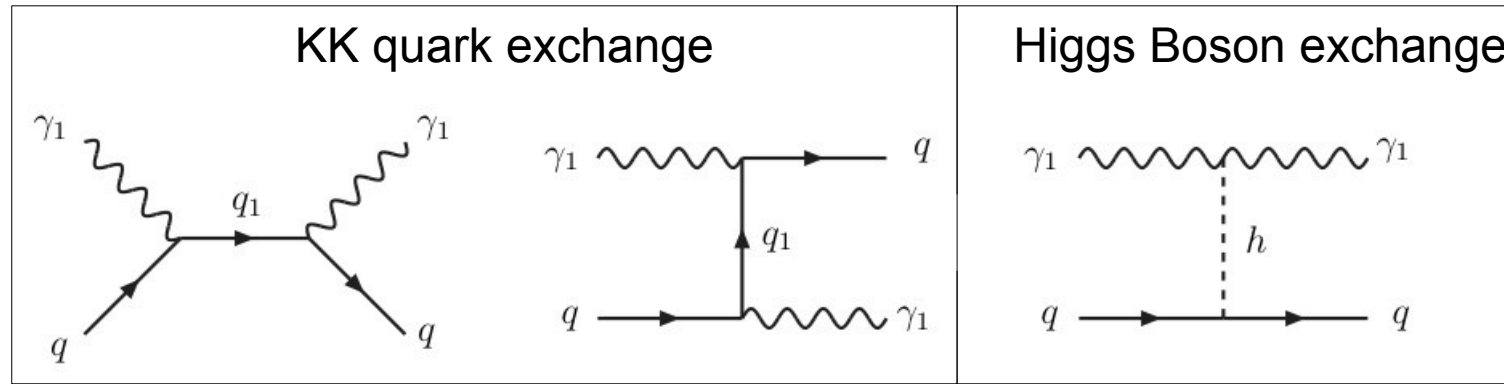
$Z_1$

- $Z_1$  and  $W_1^\pm$  are degenerate
- gluon is heavier than  $Z_1$  by 20%
- all other particles are heavier than  $Z_1$  by 10%



Computations of the relic density for 6D including coannihilations do not exist yet.

Feynman diagrams for  $\gamma_1$ -quark scattering (others are similar):



spin-dependent interaction

scalar interaction

## Important parameters

SI WIMP-nucleon couplings  $f_n, f_p$

SD WIMP-nucleon couplings  $a_n, a_p$

Higgs mass  $m_h$

degeneracy parameter  $\Delta = \frac{m_{q_1} - m_{\gamma_1}}{m_{\gamma_1}}$

$$\sigma_{\text{scalar}} = \frac{m_T^2}{4\pi(m_{\gamma_1} + m_T)^2} [Zf_p + (A-Z)f_n]^2$$

$$f_{p,n} = \sum_{u,d,s} (\beta_q + \gamma_q) \frac{m_{p,n}}{m_q} f_{T_q}^{p,n}$$

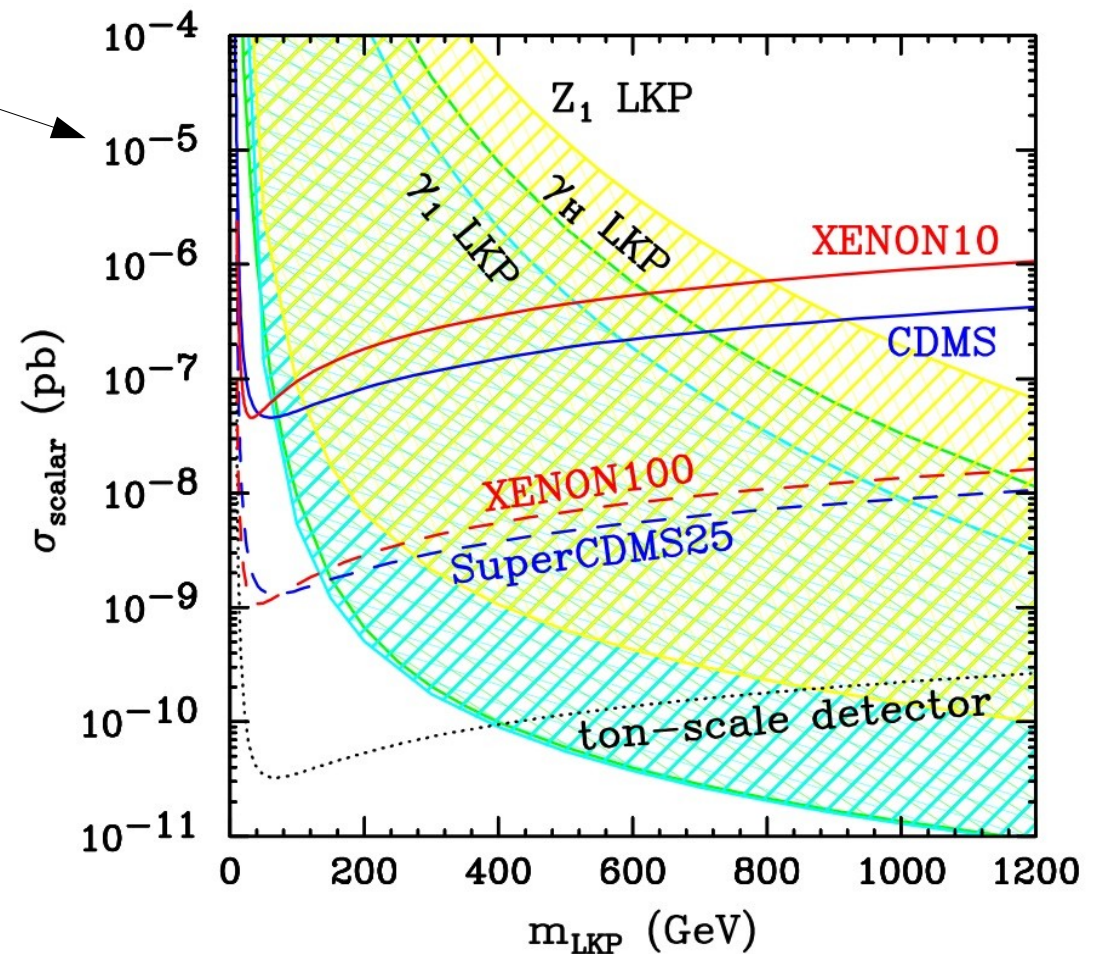
$\gamma_1$

$$\beta_q = m_q \frac{e^2}{\cos^2 \theta_W} \left[ Y_{qL}^2 \frac{m_{\gamma_1}^2 + m_{q_L}^2}{(m_{q_L}^2 - m_{\gamma_1}^2)^2} + (L \rightarrow R) \right]$$

$$\gamma_q = m_q \frac{e^2}{2 \cos^2 \theta_W} \frac{1}{m_h^2}$$

$$m_h = 120 \text{ GeV} \quad 0.01 < \Delta < 0.5$$

- significant enhancement of cross sections for small  $\Delta$
- CDMS and Xenon10 already exclude small mass splittings
- future ton-scale experiments should cover most of the interesting parameter space





$$\sigma_{spin} = \frac{32}{\pi} G_F^2 \mu^2 \frac{J_N + 1}{J_N} (a_p \langle S_p \rangle + a_n \langle S_n \rangle)^2$$

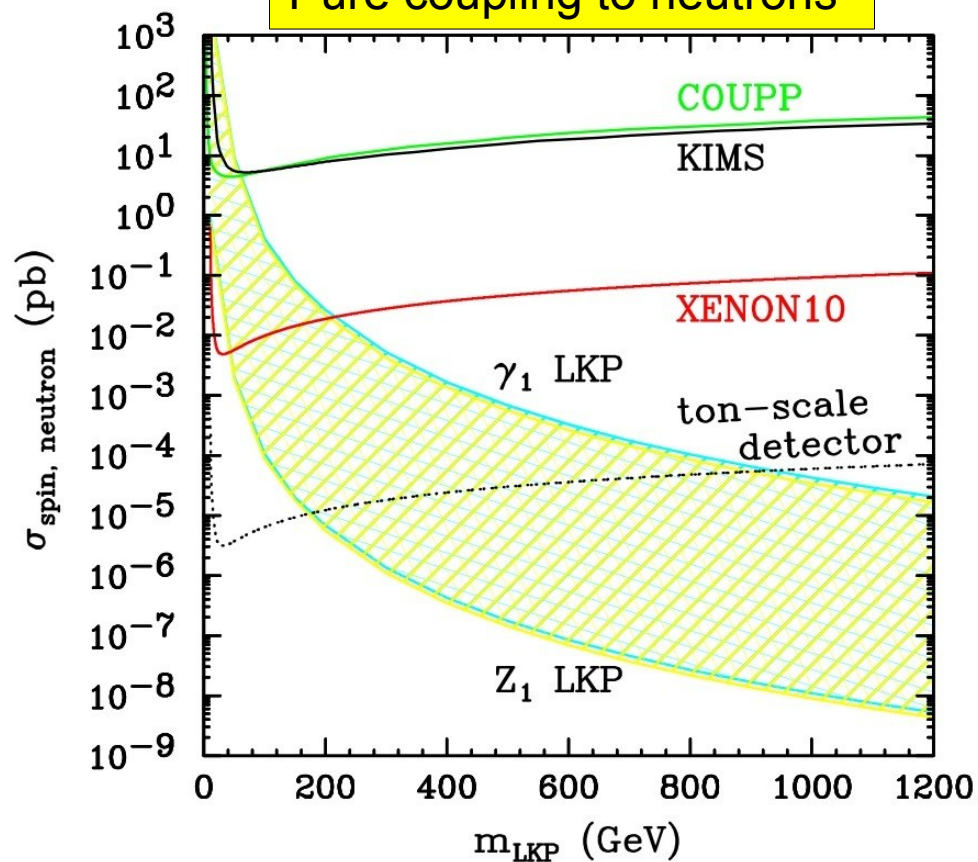
$a_{p,n}$  contain the whole theoretical model-dependence

$\gamma_1$

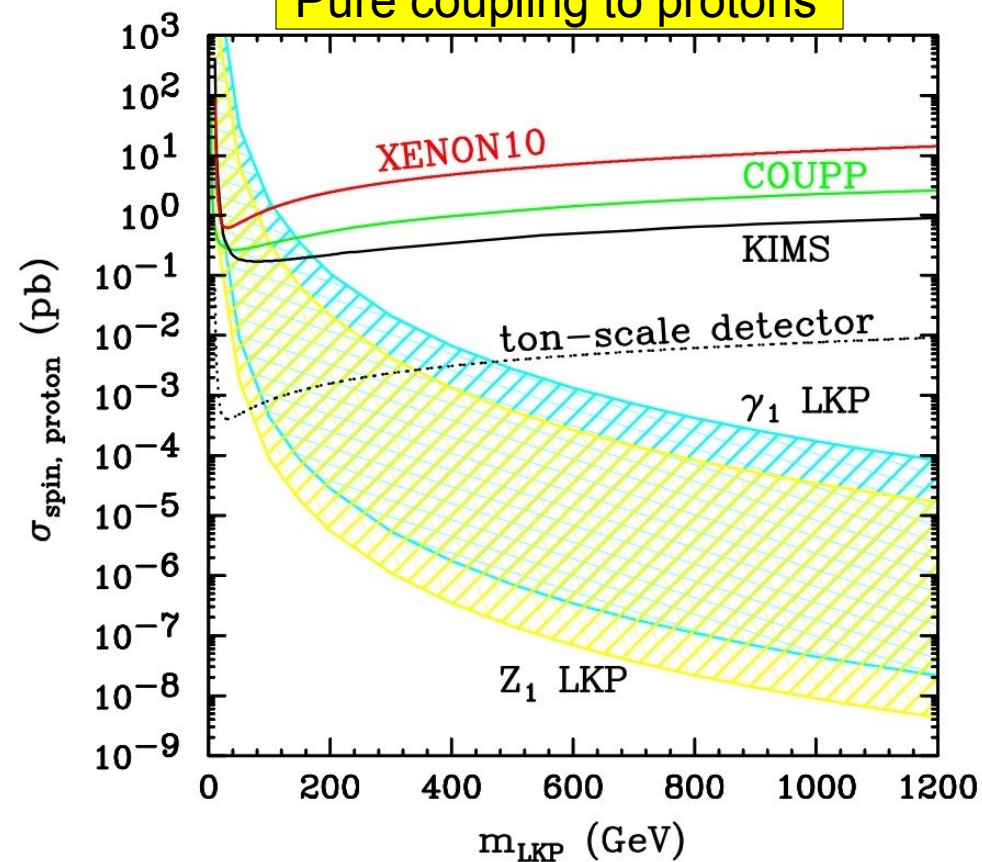
$$a_{p,n} = \frac{e^2}{4\sqrt{3} G_F \cos^2 \theta_W} \sum_{u,d,s} \left[ \frac{Y_{qL}^2}{m_{q_L}^2 - m_{\gamma_1}^2} + (L \rightarrow R) \right] \Delta_q^{p,n}$$

$$0.01 < \Delta < 0.5$$

Pure coupling to neutrons



Pure coupling to protons



Proton and neutron SD cross sections are exactly equal in the case of  $Z_1$ .

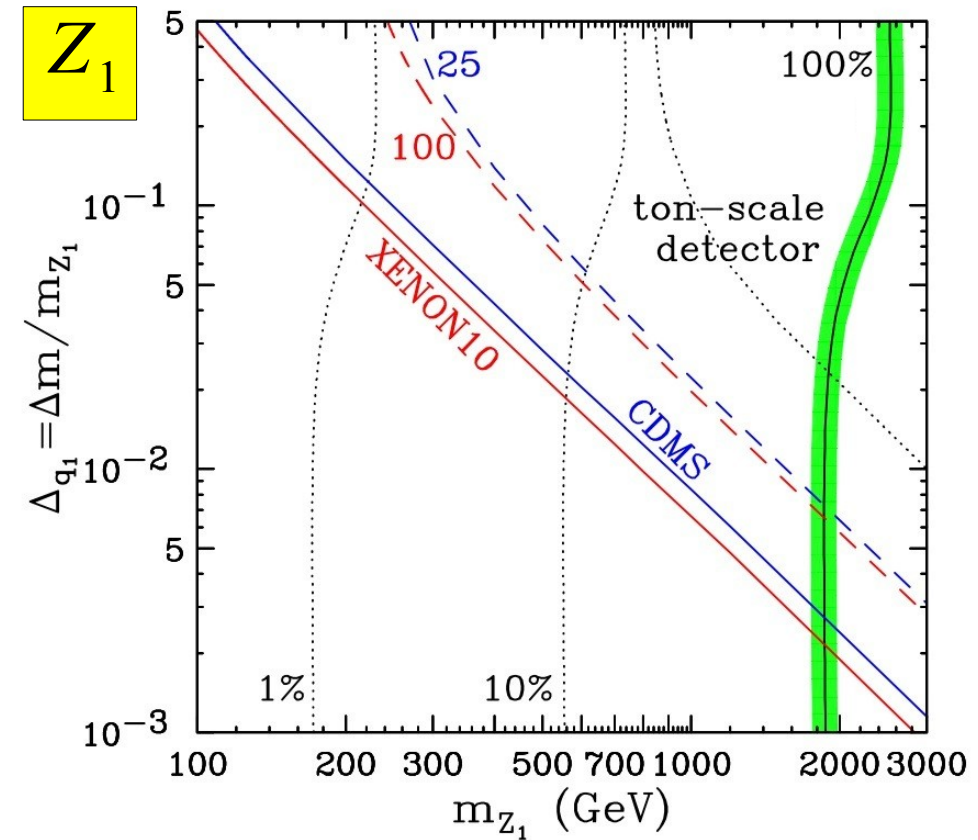
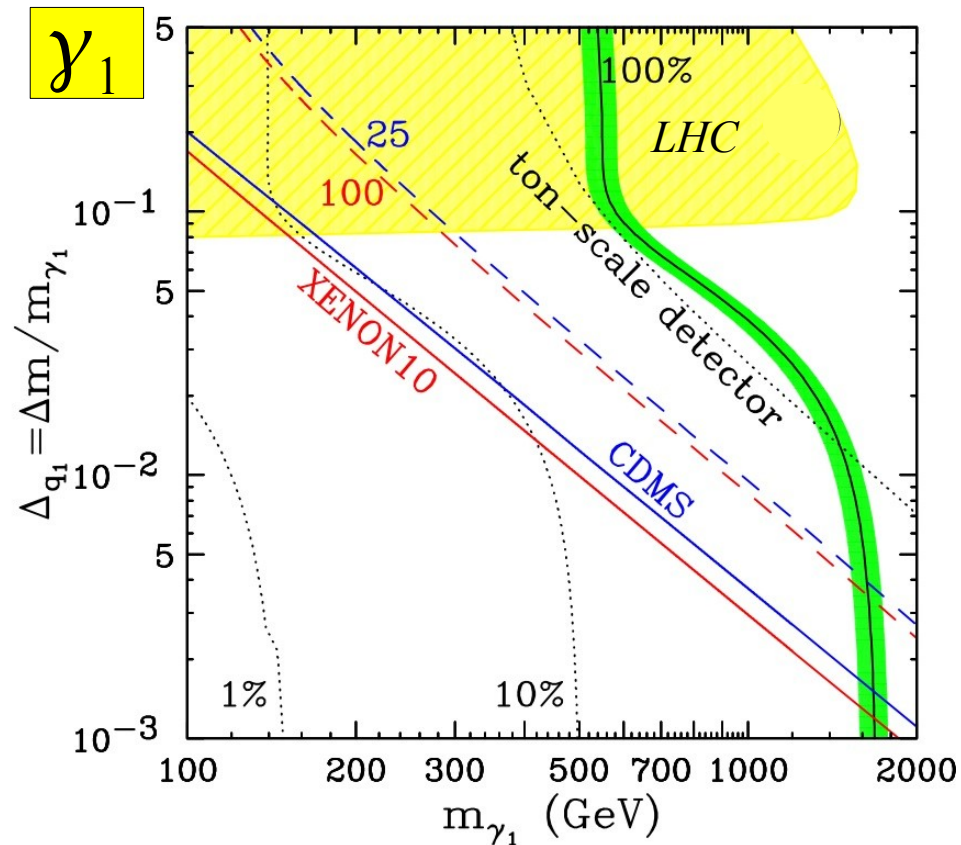
Neutron SD cross sections are approximately equal for  $\gamma_1$  and  $Z_1$ .



- free parameters: LKP mass and  $\Delta$
- Higgs mass is fixed at 120 GeV

Include....

- direct detection limits
- relic density constraints
- collider studies (hep-ph/0205314)



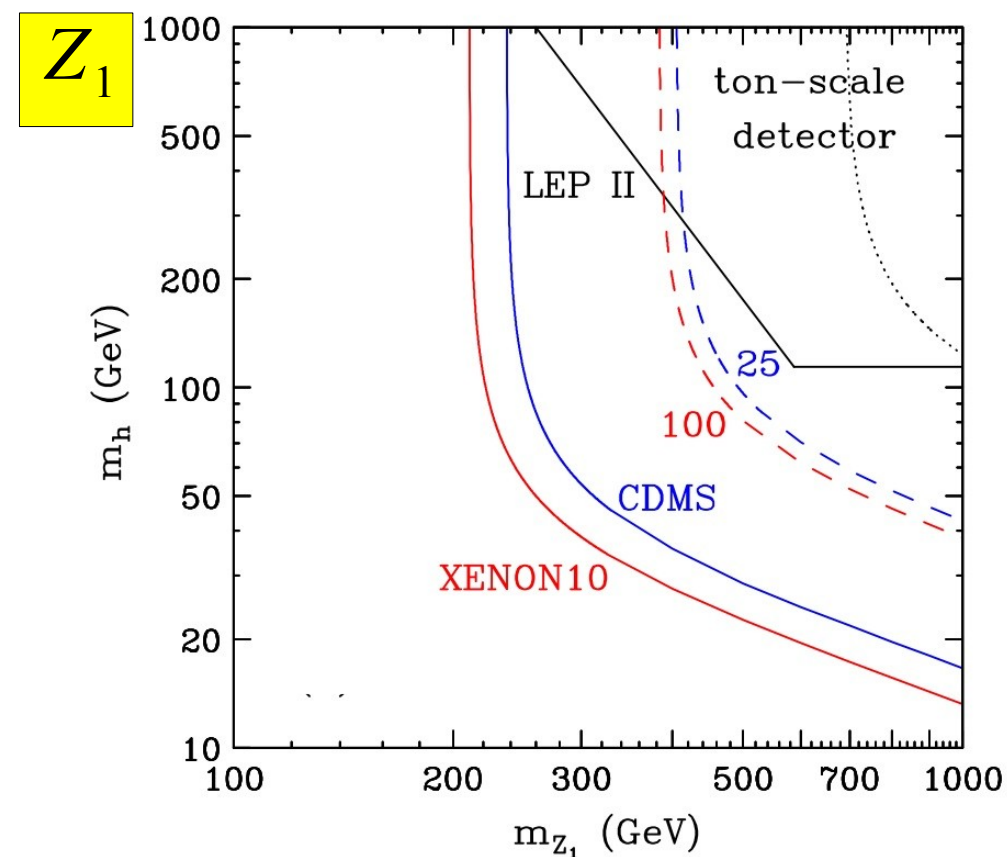
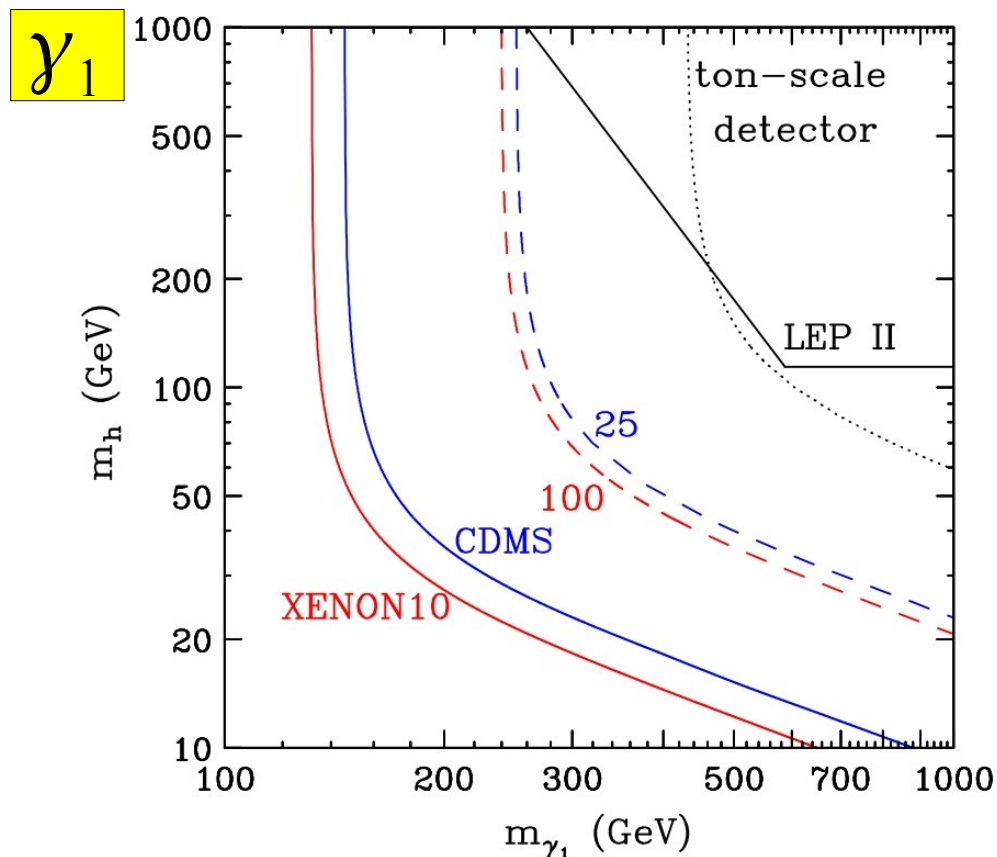
The three probes are highly complementary.

**Cosmology** provides upper limit on LKP masses.  
**Colliders** are sensitive to large  $\Delta$ .  
**Direct detection experiments** are sensitive to small  $\Delta$ .

- free parameters: LKP mass and  $m_h$
- Fix  $\Delta$  at 0.1

Include....

- direct detection limits
- collider studies (hep-ph/0605207)

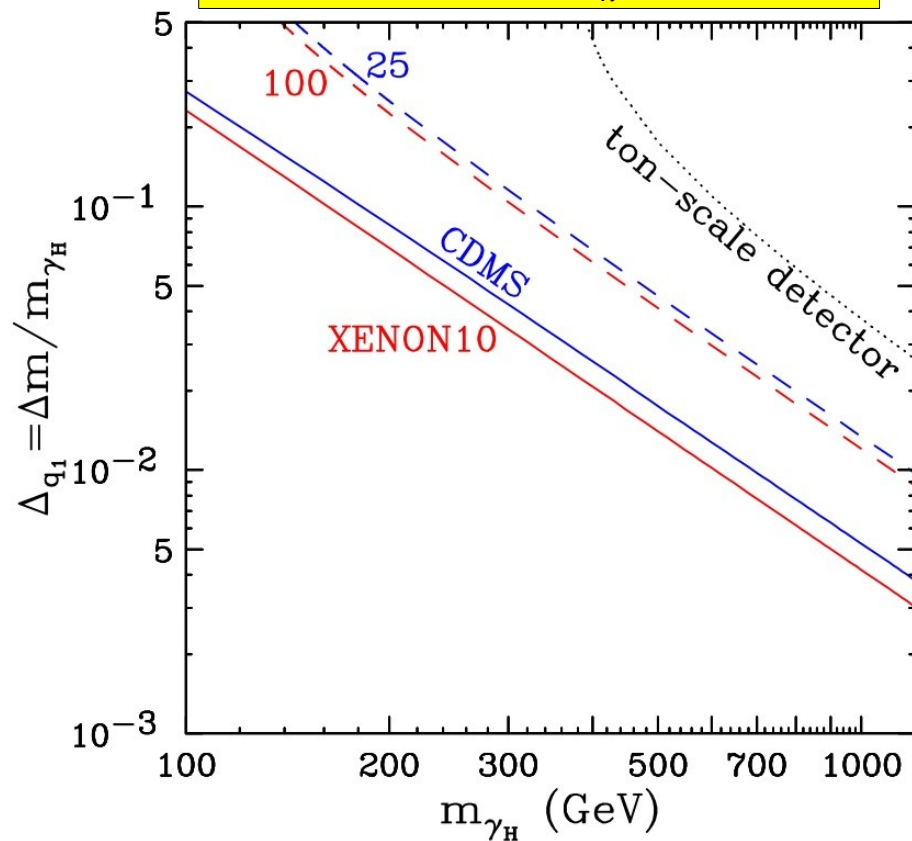


LKP mass and  $\Delta$  are primary parameters.  $m_h$  plays only a secondary role.

Future direct detection experiments only probe a small part of the parameter space.

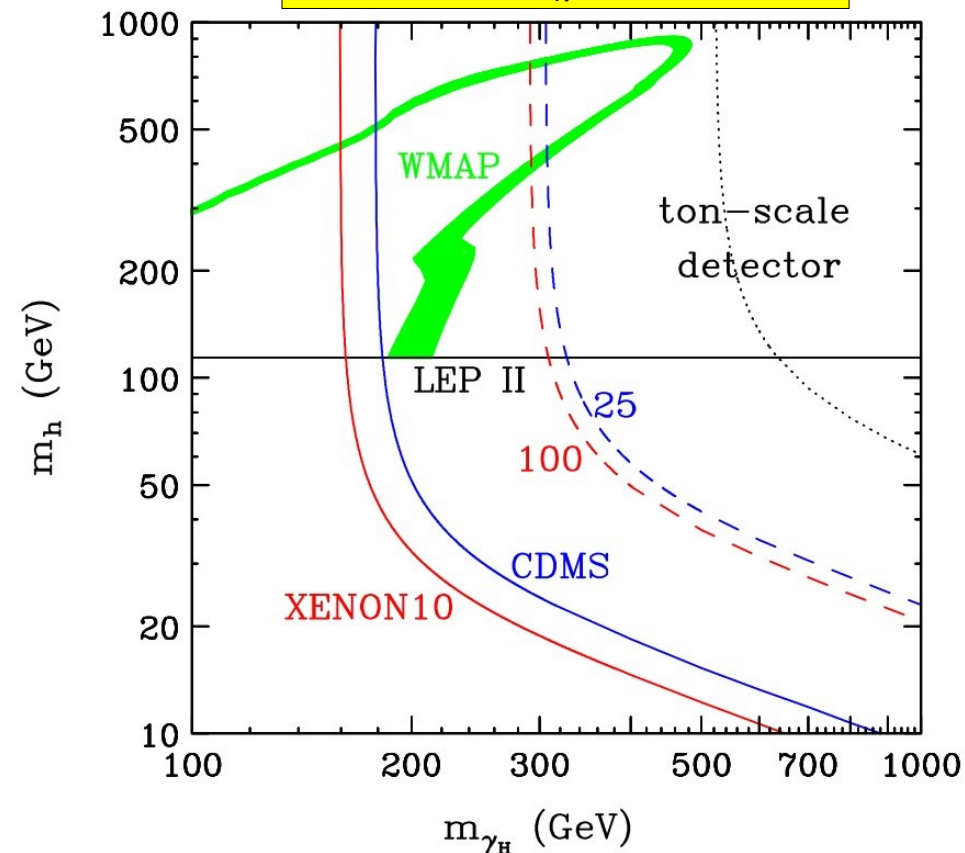
LHC will be able to test the whole parameter space shown here.

Limits on  $\Delta$  for  $m_h = 120 \text{ GeV}$



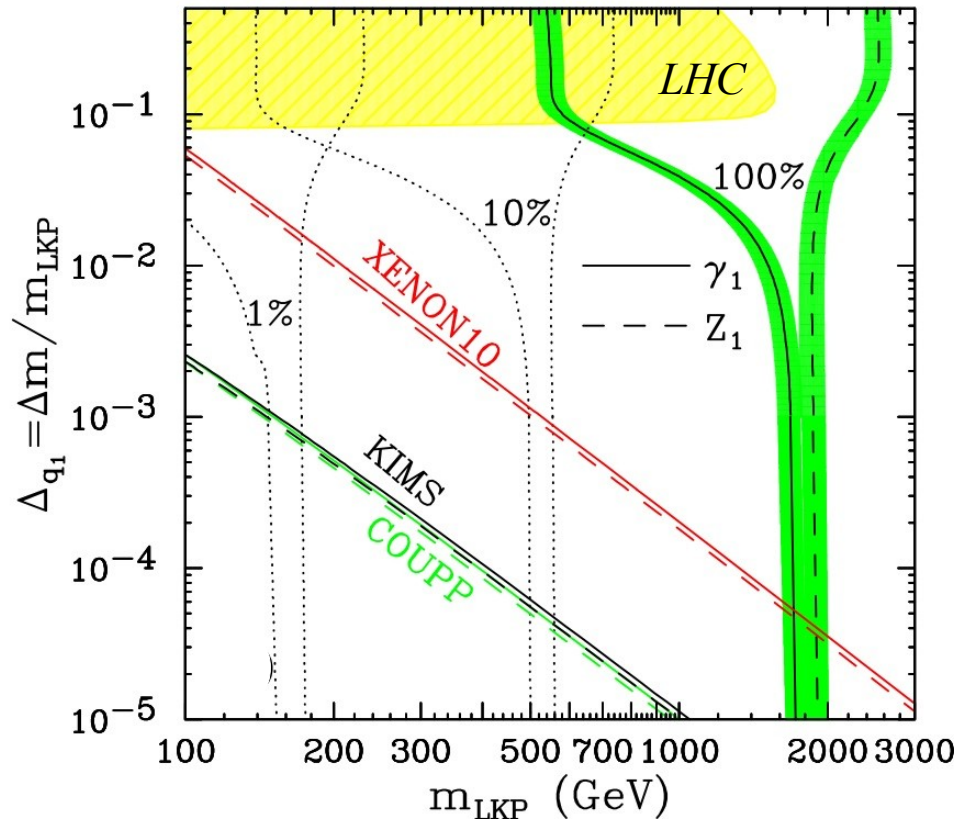
$\gamma_H$

Limits on  $m_h$  for  $\Delta = 0.1$

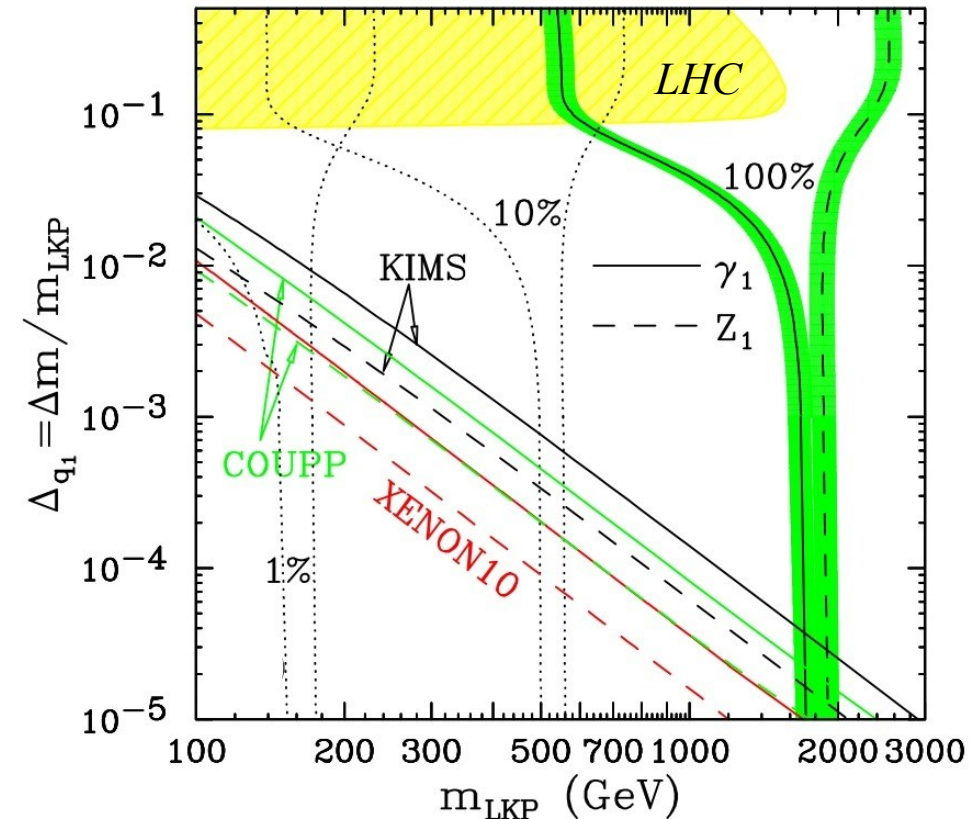


Limits on  $\Delta$  can also be computed considering spin-dependent interactions.

Pure coupling to neutrons



Pure coupling to protons

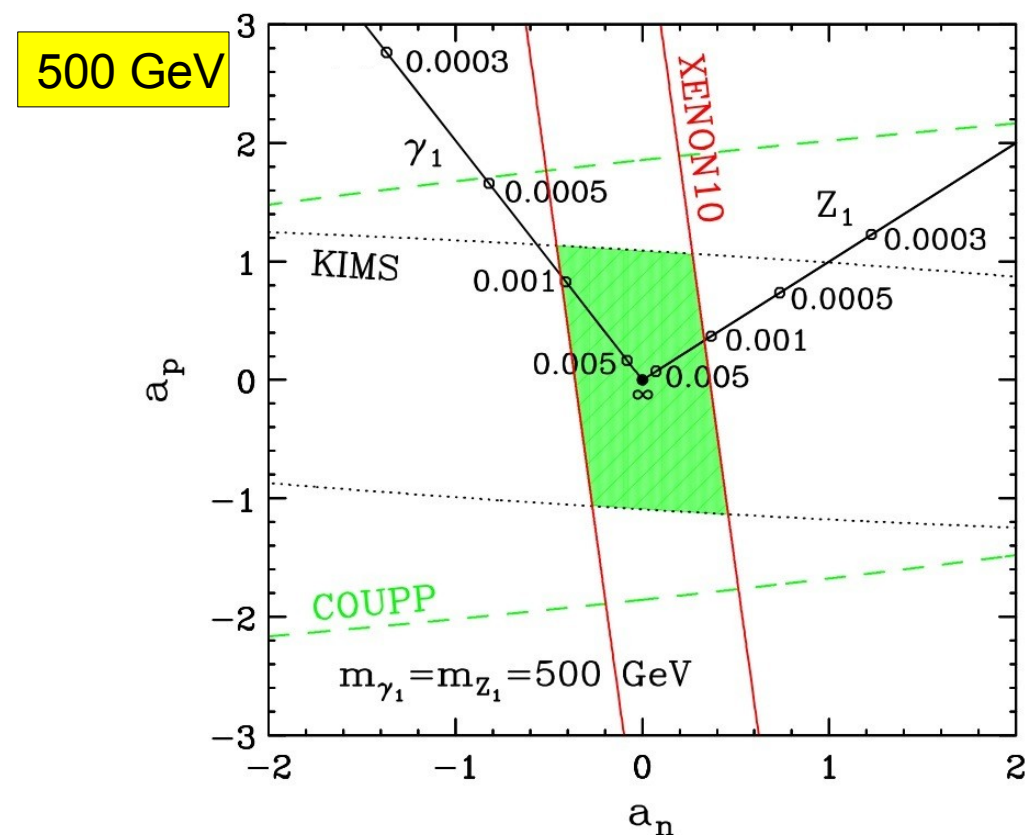
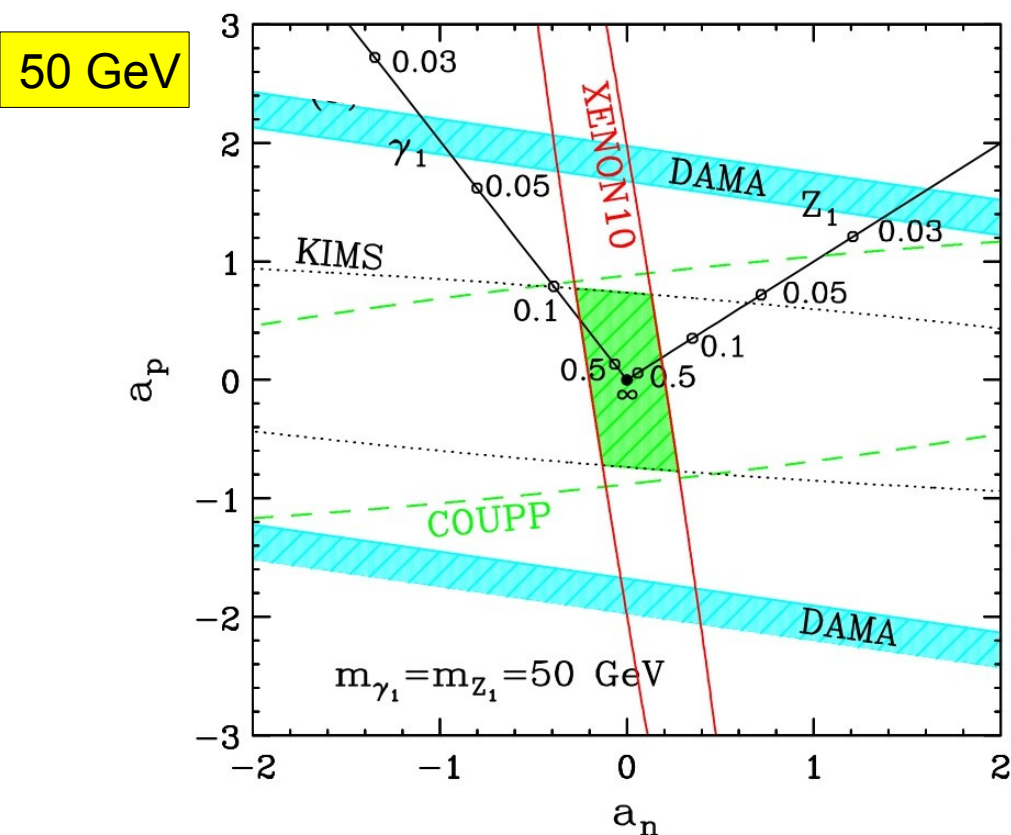


SD constraints are about an order of magnitude smaller than the SI limits.

The experiments' sensitivities to both interactions crucially depend on the used target material.



- free parameters  $a_p$  and  $a_n$
- limits from Xenon10: Introduce polar coordinates in  $a_p - a_n$  plane.  
 → Scan over  $\theta$ .



Combining limits from odd-neutron and odd-proton experiments substantially diminishes the allowed parameter space.

## What has been done?

Comprehensive analysis of 5D and 6D Kaluza-Klein dark matter including constraints from...

- direct detection experiments
- collider studies
- cosmology

## Results

- All three approaches are complementary and have the potential to cover a huge part of the relevant parameter space.

Direct detection experiments restrict small values of  $\Delta$ .

Colliders are sensitive to large  $\Delta$ s.

Cosmology rules out large LKP masses.

- Reasonable parameters to explore the KK phenomenology are  $\Delta$  and  $m_{\text{LKP}}$ .
- Coannihilation processes are of crucial importance for relic density calculations.

## What is missing?

- detailed LHC studies for small  $\Delta$
- further relic density computations for e.g. the  $\mathcal{Y}_H$  including coannihilations